



PUBLISHED FOR SISSA BY SPRINGER

RECEIVED: January 21, 2015

ACCEPTED: January 22, 2015

PUBLISHED: February 11, 2015

# Erratum: Improved measurements of the neutrino mixing angle $\theta_{13}$ with the Double Chooz detector



## The Double Chooz collaboration

Y. Abe,<sup>aa</sup> J.C. dos Anjos,<sup>e</sup> J.C. Barriere,<sup>n</sup> E. Baussan,<sup>v</sup> I. Bekman,<sup>a</sup> M. Bergevin,<sup>i</sup> T.J.C. Bezerra,<sup>y</sup> L. Bezrukov,<sup>m</sup> E. Blucher,<sup>f</sup> C. Buck,<sup>s</sup> J. Busenitz,<sup>b</sup> A. Cabrera,<sup>d</sup> E. Caden,<sup>j</sup> L. Camilleri,<sup>h</sup> R. Carr,<sup>h</sup> M. Cerrada,<sup>g</sup> P.-J. Chang,<sup>o</sup> E. Chauveau,<sup>y</sup> P. Chimenti,<sup>ae</sup> A.P. Collin,<sup>s</sup> E. Conover,<sup>f</sup> J.M. Conrad,<sup>r</sup> J.I. Crespo-Anadón,<sup>g</sup> K. Crum,<sup>f</sup> A.S. Cucoanes,<sup>w</sup> E. Damon,<sup>j</sup> J.V. Dawson,<sup>d</sup> J. Dhooghe,<sup>i</sup> D. Dietrich,<sup>ad</sup> Z. Djurcic,<sup>c</sup> M. Dracos,<sup>v</sup> M. Elnimr,<sup>b</sup> A. Etenko,<sup>q</sup> M. Fallot,<sup>w</sup> F. von Feilitzsch,<sup>ac</sup> J. Felde,<sup>i,1</sup> S.M. Fernandes,<sup>b</sup> V. Fischer,<sup>n</sup> D. Franco,<sup>d</sup> M. Franke,<sup>ac</sup> H. Furuta,<sup>y</sup> I. Gil-Botella,<sup>g</sup> L. Giot,<sup>w</sup> M. Göger-Neff,<sup>ac</sup> L.F.G. Gonzalez,<sup>af</sup> L. Goodenough,<sup>c</sup> M.C. Goodman,<sup>c</sup> C. Grant,<sup>i</sup> N. Haag,<sup>ac</sup> T. Hara,<sup>p</sup> J. Haser,<sup>s</sup> M. Hofmann,<sup>ac</sup> G.A. Horton-Smith,<sup>o</sup> A. Hourlier,<sup>d</sup> M. Ishitsuka,<sup>aa</sup> J. Jochum,<sup>ad</sup> C. Jollet,<sup>v</sup> F. Kaether,<sup>s</sup> L.N. Kalousis,<sup>ag</sup> Y. Kamyshev,<sup>x</sup> D.M. Kaplan,<sup>l</sup> T. Kawasaki,<sup>t</sup> E. Kemp,<sup>af</sup> H. de Kerret,<sup>d</sup> D. Kryn,<sup>d</sup> M. Kuze,<sup>aa,2</sup> T. Lachenmaier,<sup>ad</sup> C.E. Lane,<sup>j</sup> T. Lasserre,<sup>n,d</sup> A. Letourneau,<sup>n</sup> D. Lhuillier,<sup>n</sup> H.P. Lima Jr,<sup>e</sup> M. Lindner,<sup>s</sup> J.M. López-Castaño,<sup>g</sup> J.M. LoSecco,<sup>u</sup> B. Lubsandorzhiev,<sup>m</sup> S. Lucht,<sup>a</sup> J. Maeda,<sup>ab,3</sup> C. Mariani,<sup>ag</sup> J. Maricic,<sup>j,4</sup> J. Martino,<sup>w</sup> T. Matsubara,<sup>ab</sup> G. Mention,<sup>n</sup> A. Mereaglia,<sup>v</sup> T. Miletic,<sup>j</sup> R. Milincic,<sup>j,4</sup> A. Minotti,<sup>v</sup> Y. Nagasaka,<sup>k</sup> Y. Nikitenko,<sup>m</sup> P. Novella,<sup>d</sup> L. Oberauer,<sup>ac</sup> M. Obolensky,<sup>d</sup> A. Onillon,<sup>w</sup> A. Osborn,<sup>x</sup> C. Palomares,<sup>g</sup> I.M. Pepe,<sup>e</sup> S. Perasso,<sup>d</sup> P. Pfahler,<sup>ac</sup> A. Porta,<sup>w</sup> G. Pronost,<sup>w</sup> J. Reichenbacher,<sup>b</sup> B. Reinhold,<sup>s,4</sup> M. Röhling,<sup>ad</sup> R. Roncin,<sup>d</sup> S. Roth,<sup>a</sup> B. Rybolt,<sup>x</sup> Y. Sakamoto,<sup>z</sup> R. Santorelli,<sup>g</sup> A.C. Schilithz,<sup>e</sup> S. Schönert,<sup>ac</sup> S. Schoppmann,<sup>a</sup> M.H. Shaevitz,<sup>h</sup> R. Sharankova,<sup>aa</sup> S. Shimojima,<sup>ab</sup> D. Shrestha,<sup>o</sup> V. Sibille,<sup>n</sup> V. Sinev,<sup>m</sup> M. Skorokhvatov,<sup>q</sup> E. Smith,<sup>j</sup> J. Spitz,<sup>r</sup> A. Stahl,<sup>a</sup> I. Stancu,<sup>b</sup> L.F.F. Stokes,<sup>ad</sup> M. Strait,<sup>f</sup> A. Stüken,<sup>a</sup> F. Suekane,<sup>y</sup> S. Sukhotin,<sup>q</sup> T. Sumiyoshi,<sup>ab</sup> Y. Sun,<sup>b,4</sup> R. Svoboda,<sup>i</sup> K. Terao,<sup>r</sup> A. Tonazzo,<sup>d</sup> H.H. Trinh Thi,<sup>ac</sup> G. Valdivieso,<sup>e</sup> N. Vassilopoulos,<sup>v</sup> C. Veyssiere,<sup>n</sup> M. Vivier,<sup>n</sup> S. Wagner,<sup>s</sup> N. Walsh,<sup>i</sup> H. Watanabe,<sup>s</sup> C. Wiebusch,<sup>a</sup> L. Winslow,<sup>r</sup> M. Wurm,<sup>ad,5</sup> G. Yang,<sup>c,l</sup> F. Yermia<sup>w</sup> and V. Zimmer<sup>ac</sup>

<sup>1</sup>Now at Department of Physics, University of Maryland, College Park, Maryland 20742, U.S.A.

<sup>2</sup>Corresponding author.

<sup>3</sup>Now at Department of Physics, Kobe University, Kobe, 657-8501, Japan.

<sup>4</sup>Now at Department of Physics & Astronomy, University of Hawaii at Manoa, Honolulu, Hawaii 96822, U.S.A.

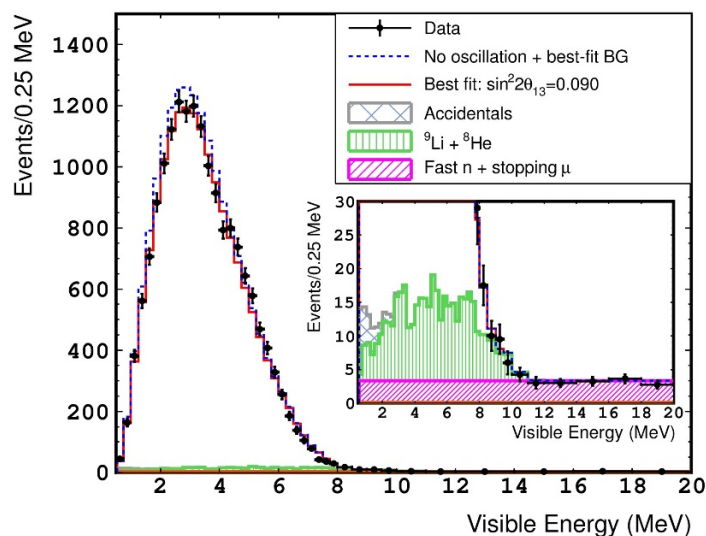
<sup>5</sup>Now at Institut für Physik and Excellence Cluster PRISMA, Johannes Gutenberg-Universität Mainz, 55128 Mainz, Germany.

- <sup>a</sup>*III. Physikalisches Institut, RWTH Aachen University, 52056 Aachen, Germany*
- <sup>b</sup>*Department of Physics and Astronomy, University of Alabama, Tuscaloosa, Alabama 35487, U.S.A.*
- <sup>c</sup>*Argonne National Laboratory, Argonne, Illinois 60439, U.S.A.*
- <sup>d</sup>*AstroParticule et Cosmologie, Université Paris Diderot, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité, 75205 Paris Cedex 13, France*
- <sup>e</sup>*Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, RJ, 22290-180, Brazil*
- <sup>f</sup>*The Enrico Fermi Institute, The University of Chicago, Chicago, Illinois 60637, U.S.A.*
- <sup>g</sup>*Centro de Investigaciones Energéticas, Medioambientales y Tecnológicas, CIEMAT, 28040, Madrid, Spain*
- <sup>h</sup>*Columbia University, New York, New York 10027, U.S.A.*
- <sup>i</sup>*University of California, Davis, California 95616, U.S.A.*
- <sup>j</sup>*Department of Physics, Drexel University, Philadelphia, Pennsylvania 19104, U.S.A.*
- <sup>k</sup>*Hiroshima Institute of Technology, Hiroshima, 731-5193, Japan*
- <sup>l</sup>*Department of Physics, Illinois Institute of Technology, Chicago, Illinois 60616, U.S.A.*
- <sup>m</sup>*Institute of Nuclear Research of the Russian Academy of Sciences, Moscow 117312, Russia*
- <sup>n</sup>*Commissariat à l'Energie Atomique et aux Energies Alternatives, Centre de Saclay, IRFU, 91191 Gif-sur-Yvette, France*
- <sup>o</sup>*Department of Physics, Kansas State University, Manhattan, Kansas 66506, U.S.A.*
- <sup>p</sup>*Department of Physics, Kobe University, Kobe, 657-8501, Japan*
- <sup>q</sup>*NRC Kurchatov Institute, 123182 Moscow, Russia*
- <sup>r</sup>*Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, U.S.A.*
- <sup>s</sup>*Max-Planck-Institut für Kernphysik, 69117 Heidelberg, Germany*
- <sup>t</sup>*Department of Physics, Niigata University, Niigata, 950-2181, Japan*
- <sup>u</sup>*University of Notre Dame, Notre Dame, Indiana 46556, U.S.A.*
- <sup>v</sup>*IPHC, Université de Strasbourg, CNRS/IN2P3, 67037 Strasbourg, France*
- <sup>w</sup>*SUBATECH, CNRS/IN2P3, Université de Nantes, Ecole des Mines de Nantes, 44307 Nantes, France*
- <sup>x</sup>*Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996, U.S.A.*
- <sup>y</sup>*Research Center for Neutrino Science, Tohoku University, Sendai 980-8578, Japan*
- <sup>z</sup>*Tohoku Gakuin University, Sendai, 981-3193, Japan*
- <sup>aa</sup>*Department of Physics, Tokyo Institute of Technology, Tokyo, 152-8551, Japan*
- <sup>ab</sup>*Department of Physics, Tokyo Metropolitan University, Tokyo, 192-0397, Japan*
- <sup>ac</sup>*Physik Department, Technische Universität München, 85748 Garching, Germany*
- <sup>ad</sup>*Kepler Center for Astro and Particle Physics, Universität Tübingen, 72076 Tübingen, Germany*
- <sup>ae</sup>*Universidade Federal do ABC, UFABC, Santo André, SP, 09210-580, Brazil*
- <sup>af</sup>*Universidade Estadual de Campinas-UNICAMP, Campinas, SP, 13083-970, Brazil*
- <sup>ag</sup>*Center for Neutrino Physics, Virginia Tech, Blacksburg, Virginia 24061, U.S.A.*

E-mail: [kuze@phys.titech.ac.jp](mailto:kuze@phys.titech.ac.jp)

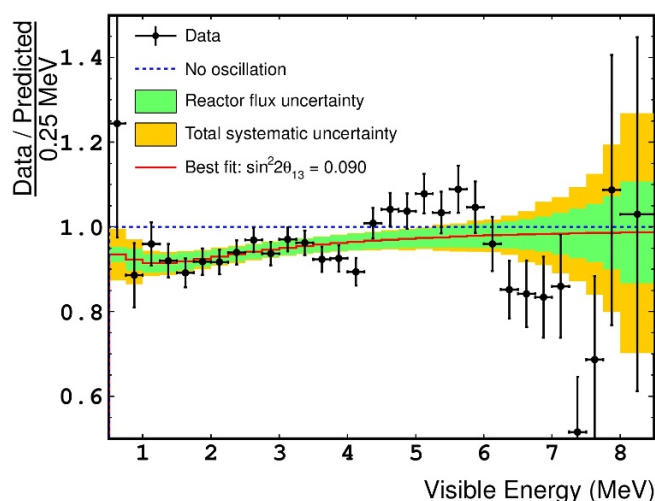
ERRATUM TO: [JHEP10\(2014\)086](#)

ARXIV EPRINT: [1406.7763](#)



**Figure 21.** The measured energy spectrum of the prompt signal (black points) superimposed on the prediction without neutrino oscillation (blue dashed line) and the best-fit with  $\sin^2 2\theta_{13} = 0.090$  (red line). Background components after the fit are also shown with different colors: accidental (grey, cross-hatched);  ${}^9\text{Li} + {}^8\text{He}$  (green, vertical-hatched); and fast neutron + stopping muons (magenta, slant-hatched).

A mistake has been found in the calculation of statistical error bars of figures 21 and 22 for bins above 8 MeV. It affects only the graphical presentation and does not change the fitted  $\theta_{13}$  value and other results of the paper. As a result of the correction, the error bars have become smaller. The figures have been replaced.



**Figure 22.** Black points show the ratio of the data, after subtraction of the background, to the non-oscillation prediction as a function of the visible energy of the prompt signal. Overlaid red line is the rate of the best-fit to the non-oscillation prediction with the reactor flux uncertainty (green) and total systematic uncertainty (orange).

**Open Access.** This article is distributed under the terms of the Creative Commons Attribution License ([CC-BY 4.0](https://creativecommons.org/licenses/by/4.0/)), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.